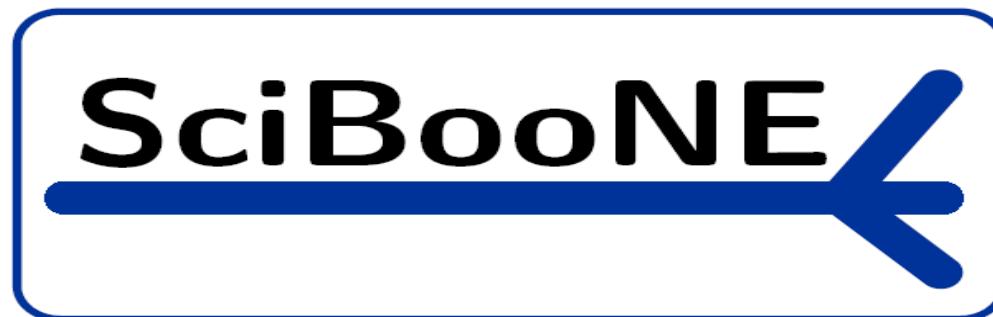


# The SciBooNE Experiment at FNAL

Yoshi Kurimoto  
Kyoto University



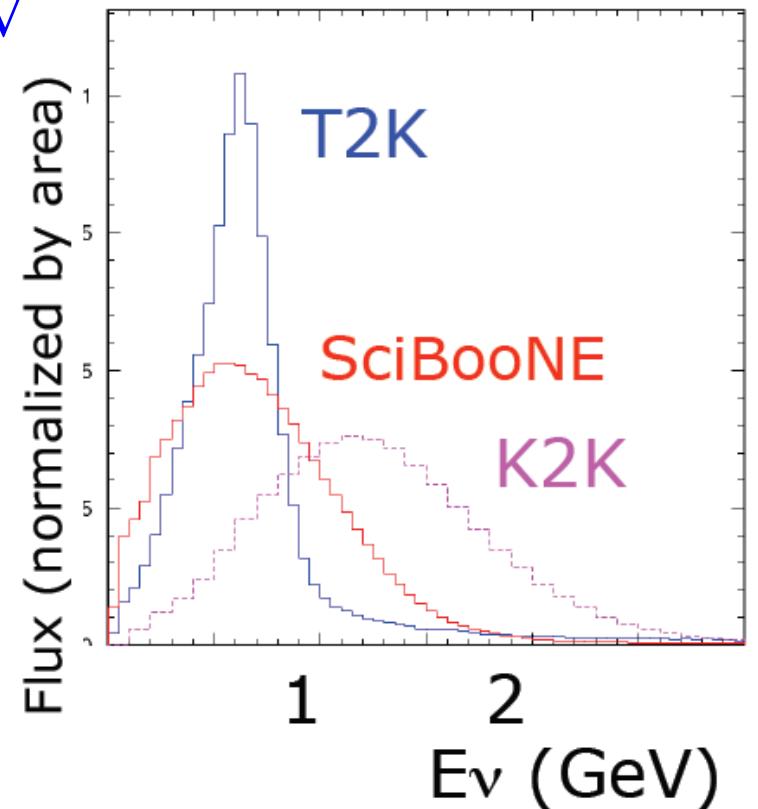
NOW08 in Italy, Sep,9.2008

# Contents

- SciBooNE experiment
- NC  $\pi^0$  measurement
- CC-coherent  $\pi$  measurement
- Summary

# SciBooNE (FNAL E954 ) Motivation

- Precise measurements of neutrino- and antineutrino-nucleus cross sections near 1 GeV
  - Essential for future neutrino oscillation experiments (T2K)
  - Few measurement in region, all low statistics
- Neutrino energy spectrum measurements and flavor measurement
  - MiniBooNE/SciBooNE joint  $\nu_\mu$  disappearance
  - $\nu_e$  constraint for MiniBooNE



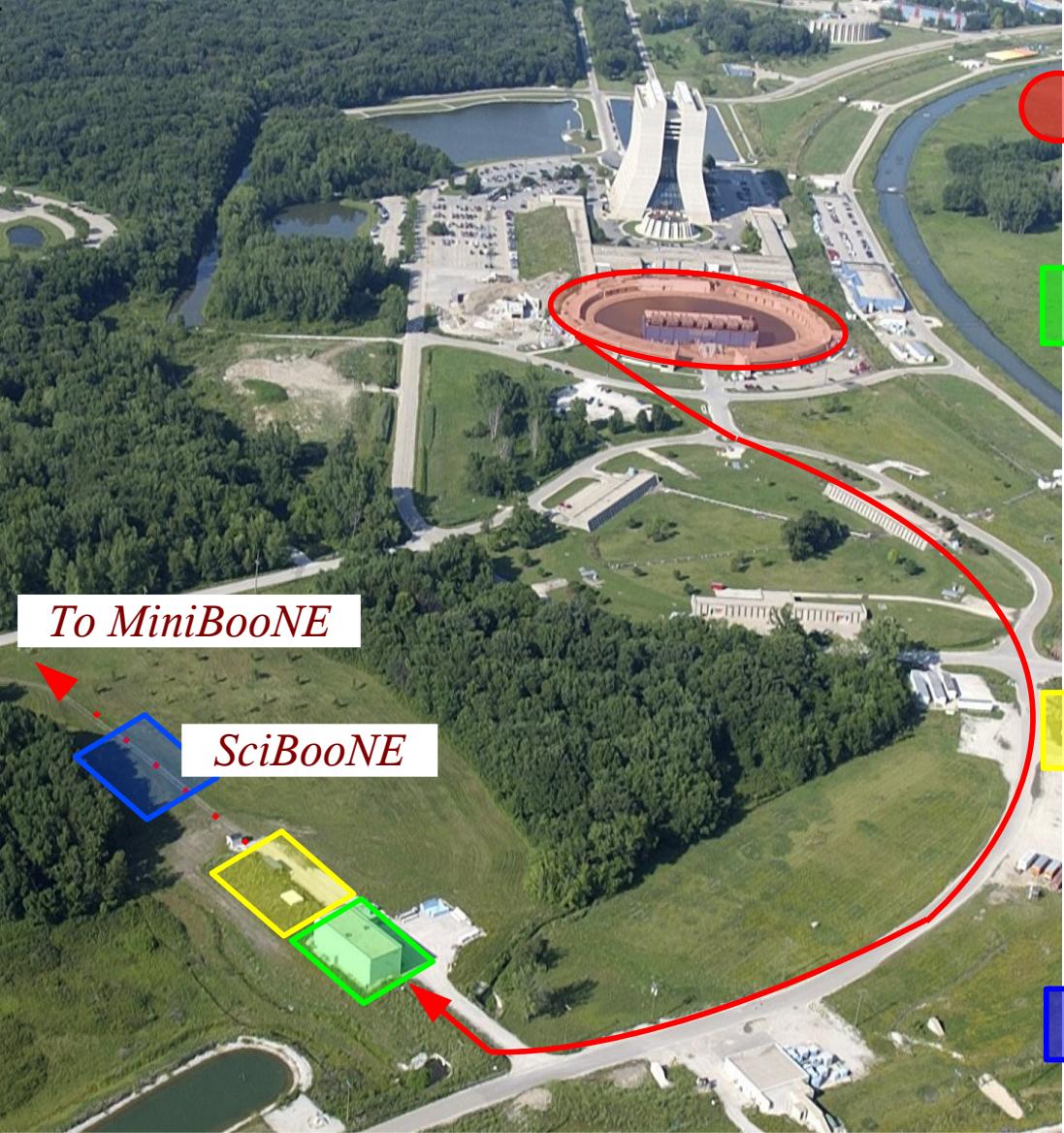
# SciBooNE Collaboration

*A selection of SciBooNE collaborators at the last collaboration meeting. March 2008*



**~70 physicists from 5 countries**

- Universitat Autònoma de Barcelona
- University of Cincinnati
- University of Colorado
- Columbia University
- Fermi National Accelerator Laboratory
- High Energy Accelerator Research Organization (KEK)
- Imperial College London\*
- Indiana University
- Institute for Cosmic Ray Research
- Kyoto University\*
- Los Alamos National Laboratory
- Louisiana State University
- Purdue University Calumet
- Università degli Studi di Roma and INFN-Roma
- Saint Mary's University of Minnesota
- Tokyo Institute of Technology
- Universidad de Valencia



## Booster Proton accelerator

- 8 GeV protons sent to target

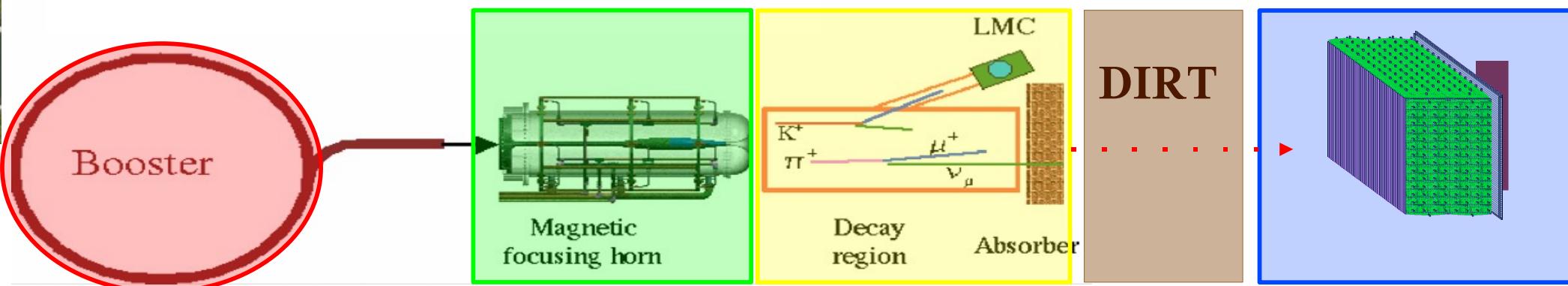
## Target Hall

- Beryllium target: 71cm long 1cm diameter
- Resultant mesons focused with magnetic horn
- Reversible horn polarity

## 50m decay volume

- Mesons decay to  $\mu$  &  $\nu_\mu$
- Short decay pipe minimises  $\mu \rightarrow \nu_e$  decay

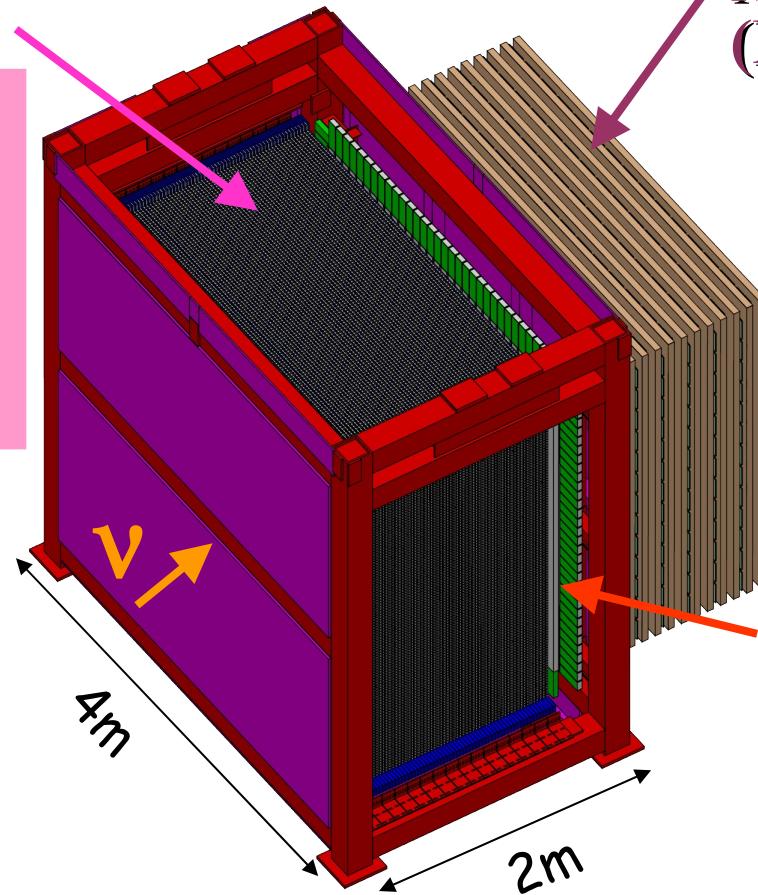
## SciBooNE located 50m from Absorber



# SciBooNE detector

**SciBar**

- 14,336 scintillator bars (15 tons)
- detect all charged particles
- $p/\pi$  separation using  $dE/dx$



**Muon Range Detector (MRD)**

- 12 2"-thick steel + scintillator planes
- measure muon momentum with range up to 1.2 GeV/c

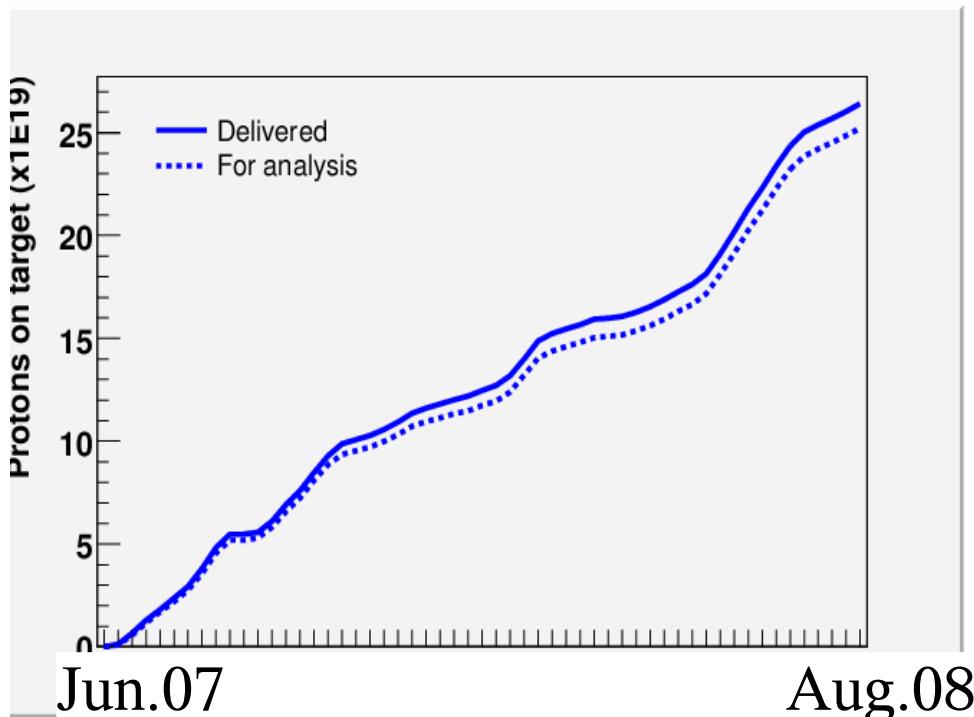
**Electron Catcher (EC)**

- spaghetti calorimeter
- 2 planes ( $11 X_0$ )
- identify  $\pi^0$  and  $\nu_e$

# Data taking

We completed data taking at Aug. 18<sup>th</sup> 2008 !!!!!!

Protons on target



- started in June 2007
- 95.5% data efficiency

• Neutrino data :  $0.99 \times 10^{20}$  POT completed!

• Antineutrino data:  $1.53 \times 10^{20}$  POT completed!

Total :  $2.52 \times 10^{20}$



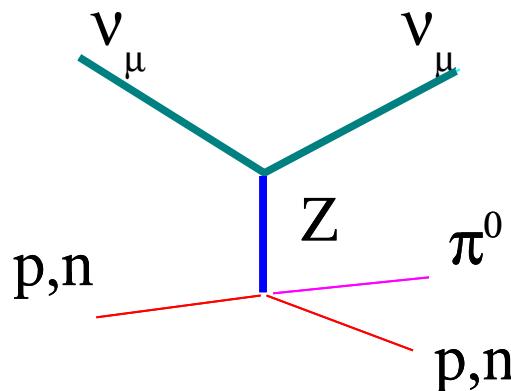
Proposal :  $2.00 \times 10^{20}$

Preliminary result from full neutrino data set are presented

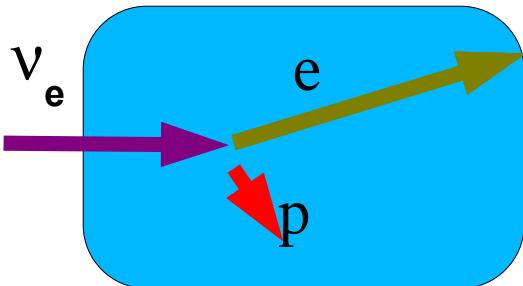
NOW08 in Italy, Sep.9.2008

# Neutral Current $\pi^0$ Production

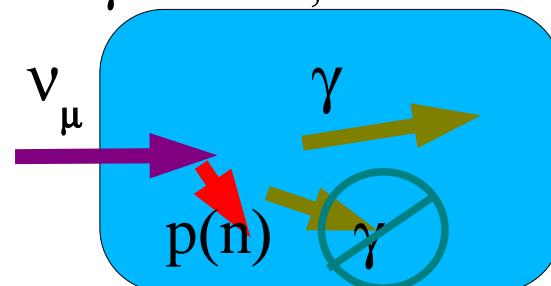
## Neutral Current $\pi^0$ ( NC $\pi^0$ )



The  $\nu_e$  signal : electron

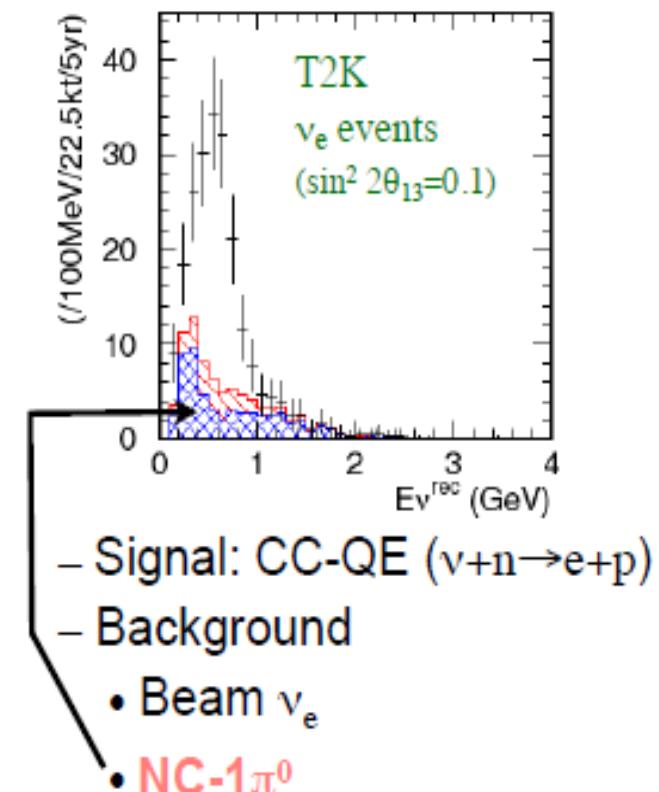


The background from NC $\pi^0$  :  
One  $\gamma$  from  $\pi^0$ , miss another  $\gamma$

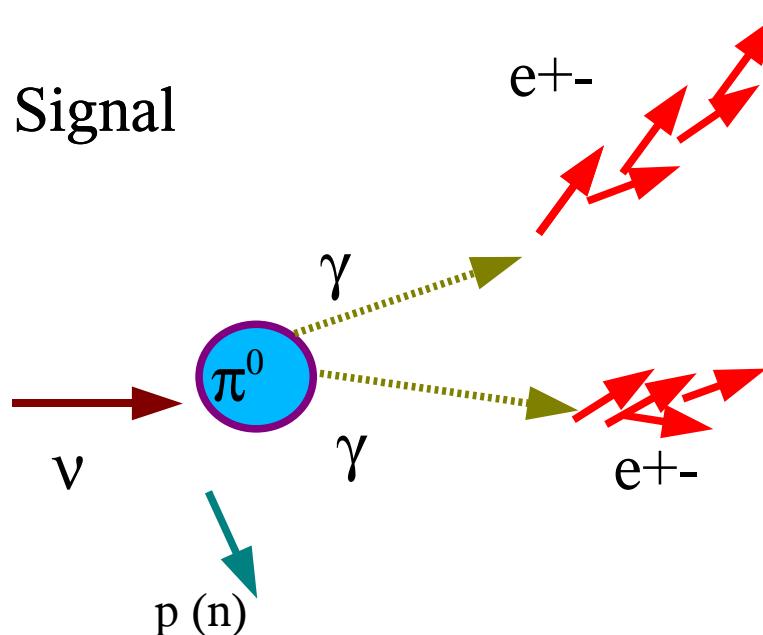


NC $\pi^0$  is the main background of the search for  $\nu_\mu$  to  $\nu_e$  oscillation

- gamma mimics electron from  $\nu_e$



# NC $\pi^0$ signal and background



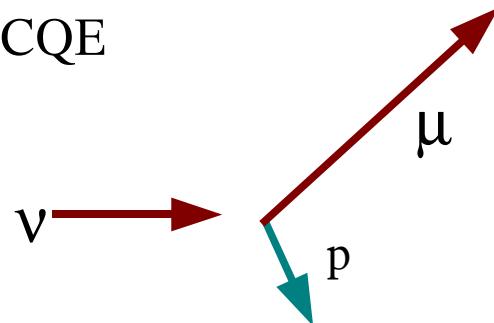
2 $\gamma$  from  $\pi^0$

- 2 tracks in Fiducial Volume
- Disconnected
- Both tracks are not  $\mu, p$

Background  $\mu, p$  common vertex or outside from detector(external)

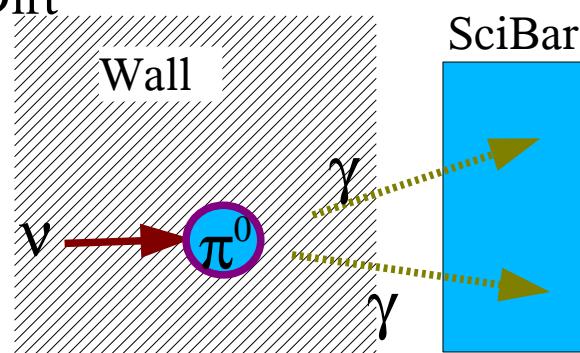
Internal B.G. :  $\nu$  int. in SciBar

ex. : CCQE



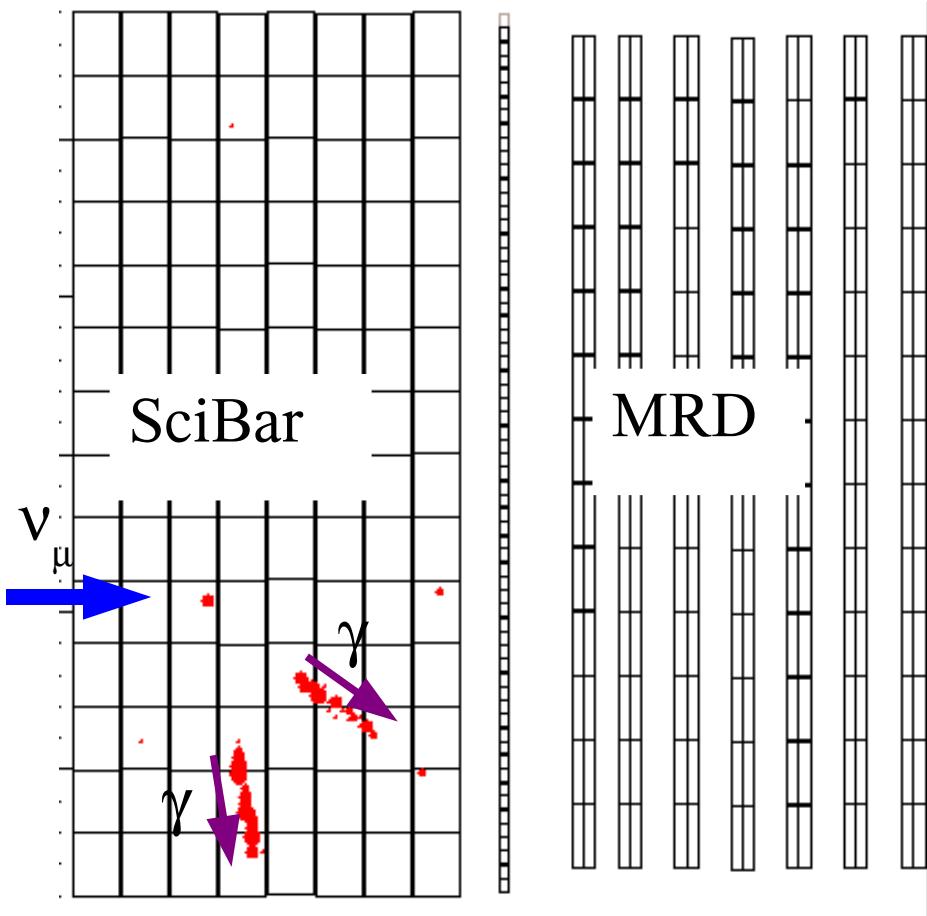
External B.G. : from outside

ex : Dirt



# Event Selection

NC $\pi^0$  Candidate



Event Selection

## 0. Pre-Selection

- At least two tracks ( $2\gamma$ )
- without 1<sup>st</sup> layer hits (reject dirt )
- Tracks Stopped in SciBar (reject  $\mu$  )

## 1. Using the track information

- Reject p using the  $dE/dX$
- Reject  $\mu$  using the decay e

## 2. Using the event topology

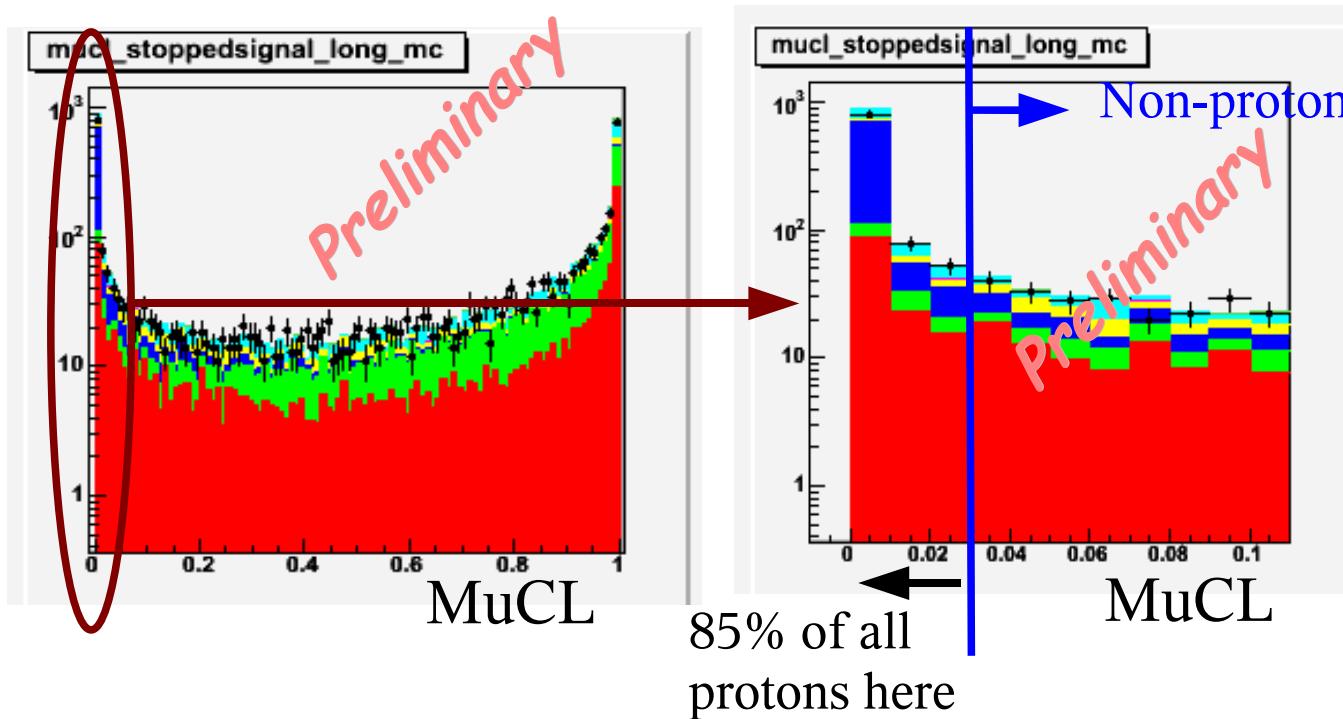
Disconnection between 2 tracks

# Proton rejection using $dE/dx$

Proton have larger energy deposit  
than any other particle

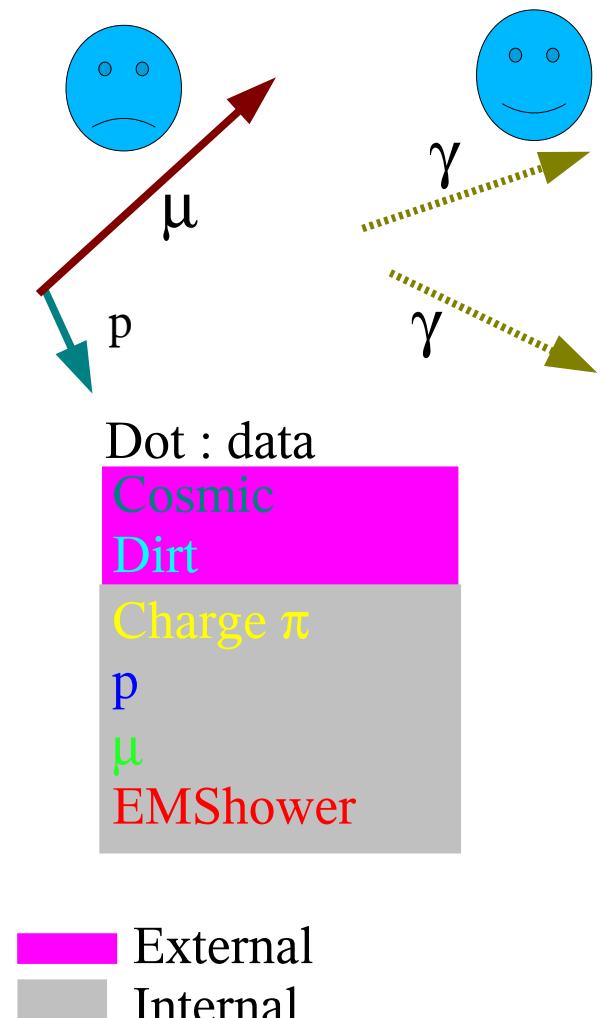


Construct "Muon Confidence Level"  
to separate proton from other particle



***Non-proton-like : MuCL > 0.03***

Events are required to have at least 2 non-proton-like tracks

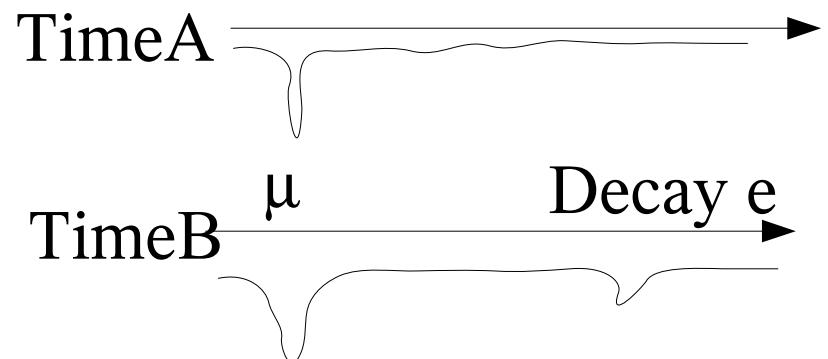
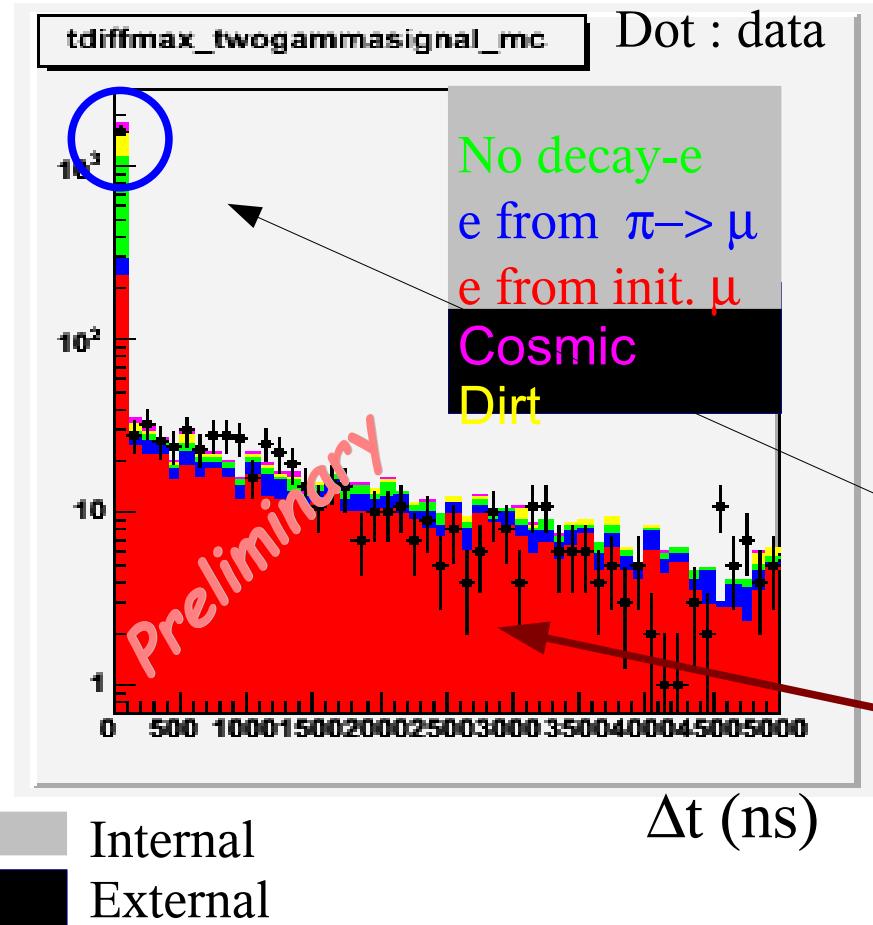


# Muon rejection with the decay electron

There are muons stopping in SciBar

These muons can be removed using the decay electron

Time difference btw track edges



$\Delta t = (\text{The latest}) \text{ TimeB} - \text{TimeA}$

Most “no decay” events is in the region  $< 100$  ns (first bin)

Reject “ $> 100$  ns”

Reject 66 % of decay-e events

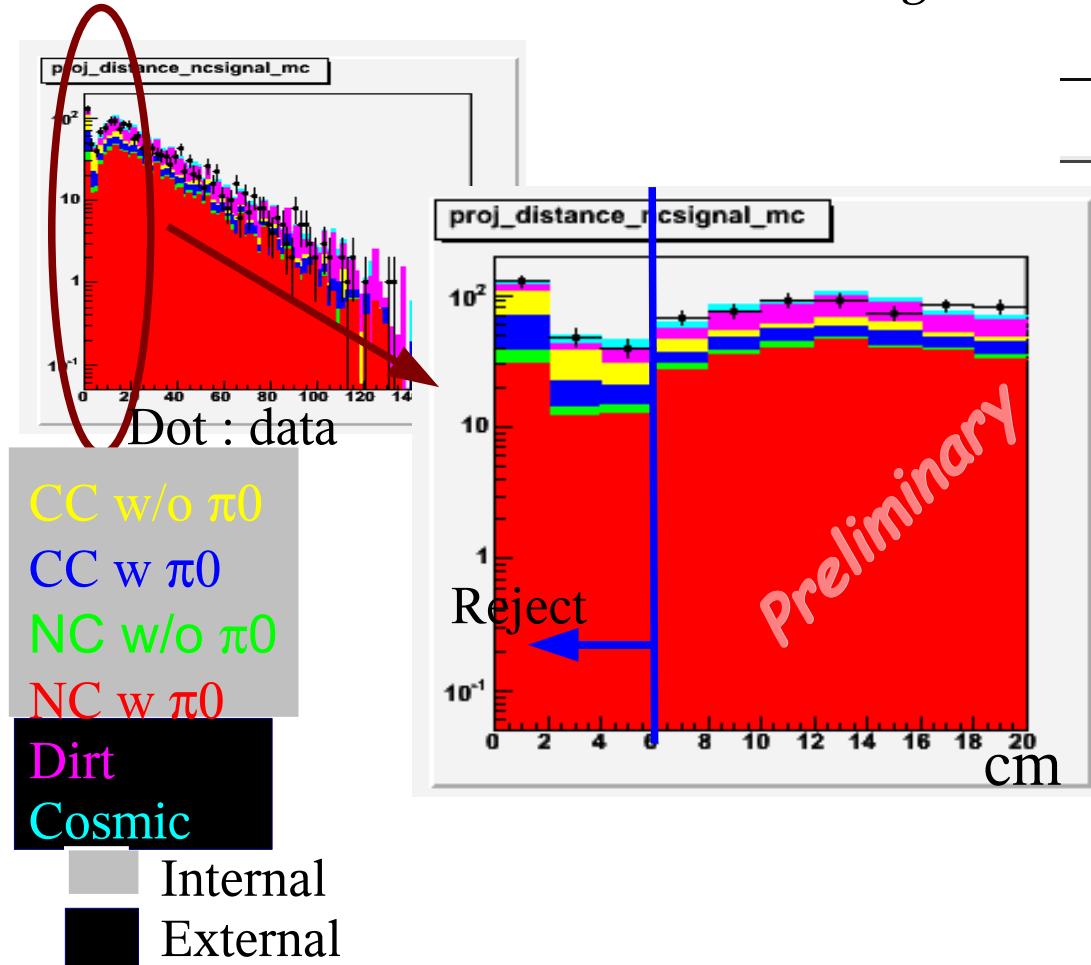
Keep 95 % of no decay-e events

<sup>12</sup>

# The disconnection between tracks

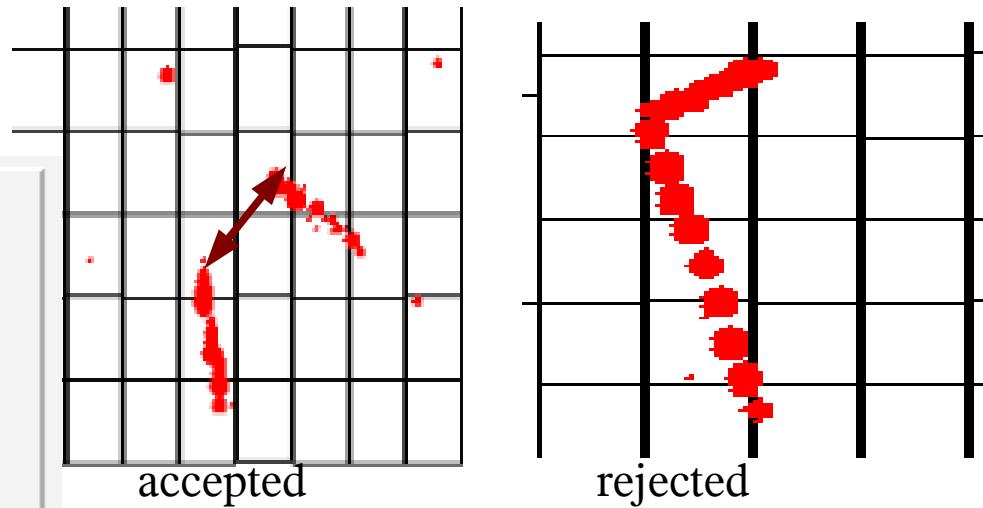
The distance between tracks (after all events selection)

Minimum 2D-distance between track edges



To get  $2\gamma$  from  $\pi^0$  and remove CC events

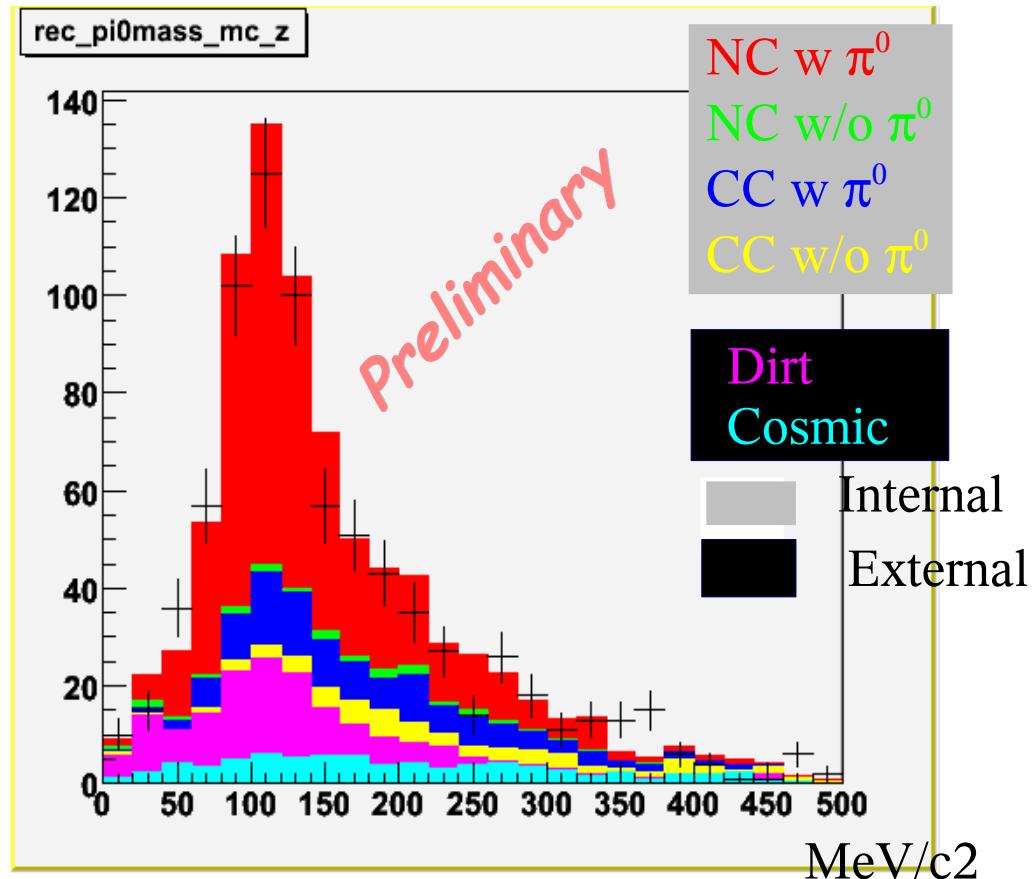
Threshold = 6cm



Reject 31 % of CC events  
Keep 92 % of  $\pi^0$  events (NC)

# $\pi^0$ reconstruction at final sample

Reconstructed Mass



- Reconstructing mass using 2 tracks in SciBar after all event selection

Clear  $\pi^0$  mass peak !  
~850 events selected  
~460  $\pi^0$  events (NC)

SciBar can  
reconstruct  $\pi^0$  !!!

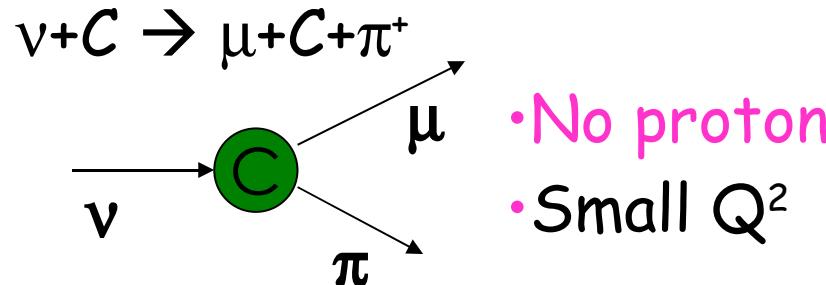


Cross section measurement

# CC-coherent $\pi^+$ measurement

## signal

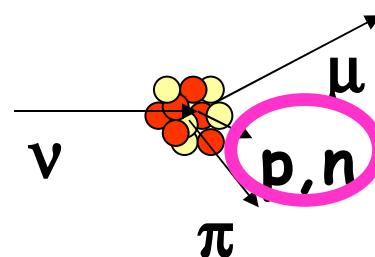
CC-coherent  $\pi$  production



## background

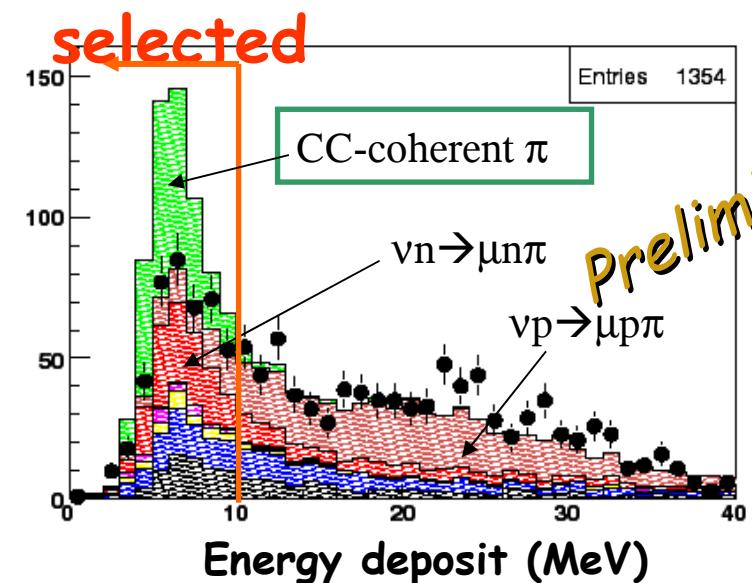
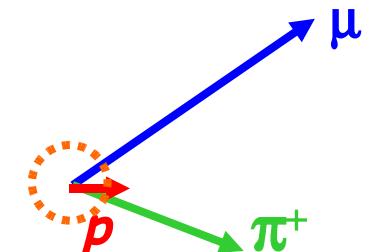
CC-resonant  $\pi$  production

- $\nu + p \rightarrow \mu + p + \pi^+$
- $\nu + n \rightarrow \mu + n + \pi^+$

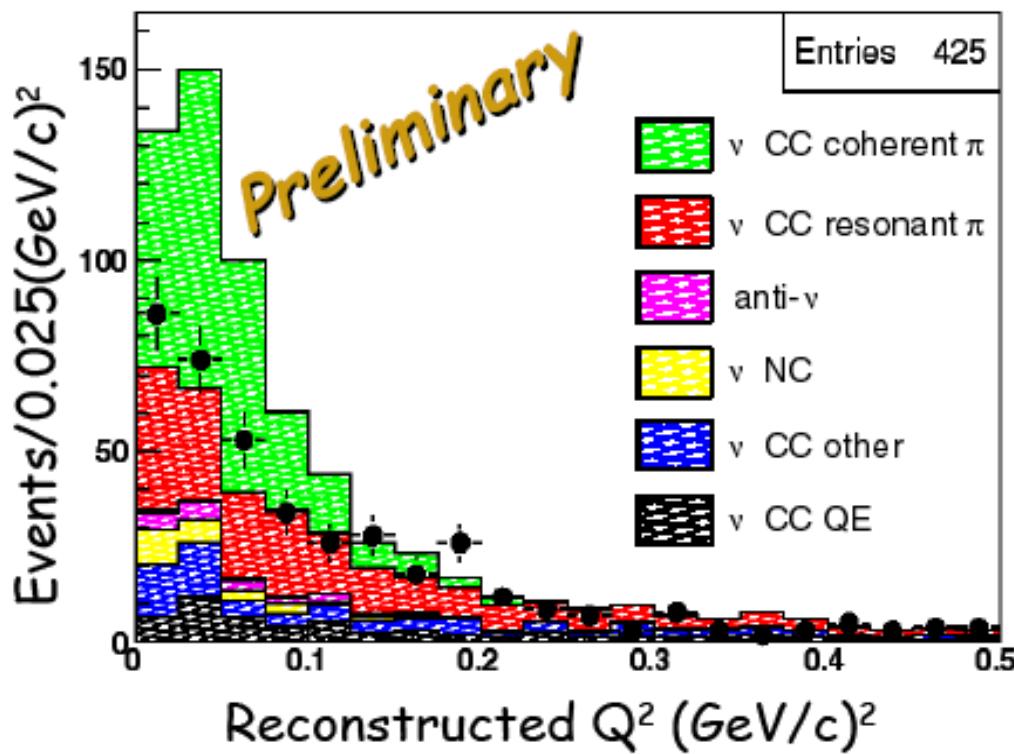


## Vertex activity

Low energy proton is identified as a large energy deposit around the vertex



# Final CC-coherent $\pi$ sample



## CC-coherent $\pi$

Efficiency 13%

Purity 40%

- \* Systematic error on background estimation is not included yet

Observed CC-coherent  $\pi$  sample in SciBooNE contains fewer events than our MC simulation, which is based on the Rein&Sehgal model

see K.Hiraide's ICHEP08 talk

# Summary

- # SciBooNE measures neutrino and antineutrino cross sections near 1 GeV
- # Data taking already was completed
- # NC  $\pi^0$  measurement
  - SciBar can reconstruct  $\pi^0$
  - Next : cross section measurement
- # Results will come soon
  - Many other analyses (CC1 $\pi$ , CCqE, CC $\pi^0$ , NCelastic,  $\nu_\mu$  disappearance,  $\nu_e$  rate ) are on-going
  - Antineutrino measurements

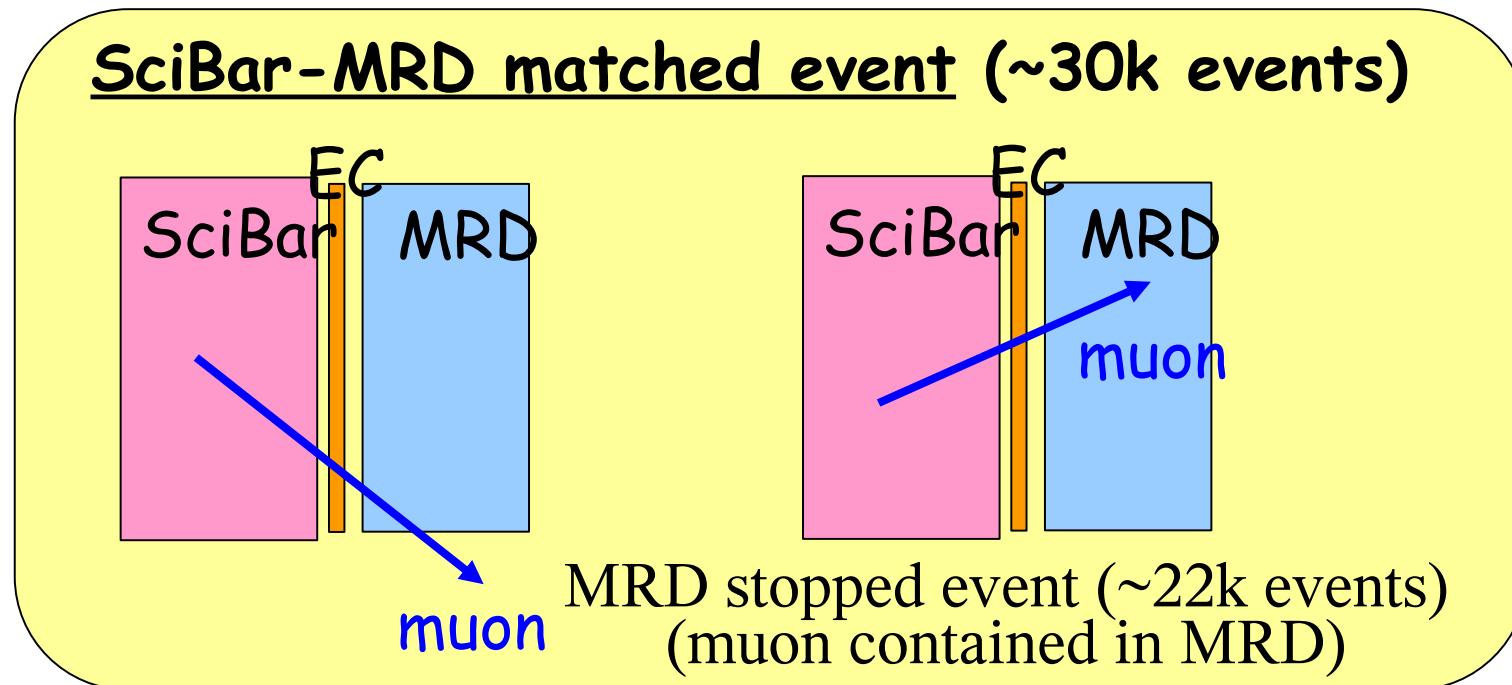
# Thank you !



Beer fight @ the day when we finished the operation !!!!

# Neutrino Spectrum Measurement

Using Charged Current Events



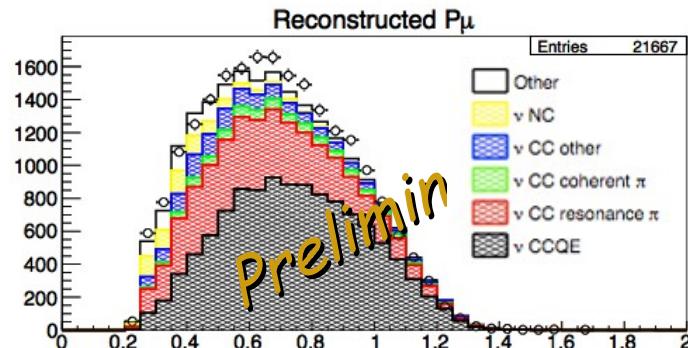
96 % pure Charged Current inclusive sample

The number of SciBar-MRD matched event is used for  
normalizing MC

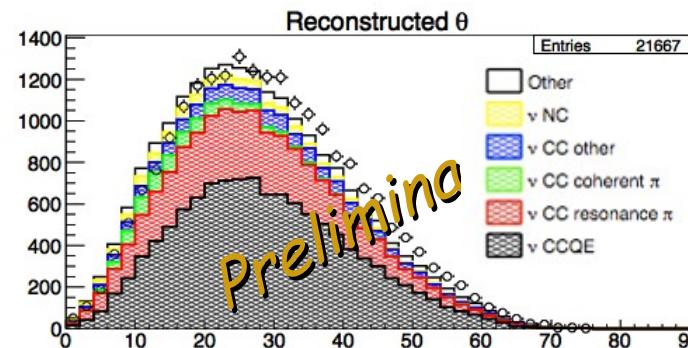
NOW08 in Italy, Sep,9.2008

# Neutrino Spectrum Measurement

Muon Kinematics of the MRD stopped sample **y.Nakajima**

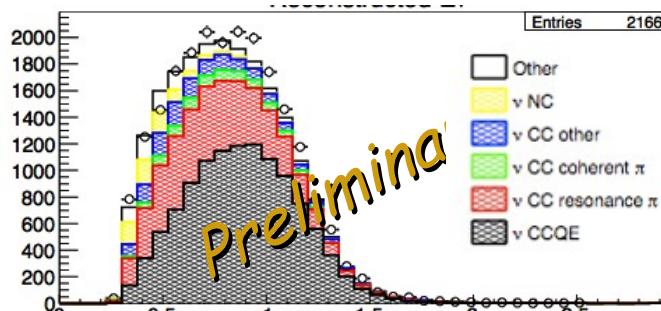


$P_\mu$  (GeV/c)



$\theta_\mu$  (degree)

Reconstructed neutrino energy assuming  
Charged Current Quasi Elastic scattering



$E_{v^{\text{rec}}}$  (GeV) NQW08 in Italy, Sep.9.2008

$$\begin{aligned}\sigma(P_\mu) &\sim 50 \text{ MeV}/c \\ \sigma(\theta_\mu) &\sim 0.9 \text{ degree} \\ M_A &= 1.1 \text{ GeV}/c^2\end{aligned}$$

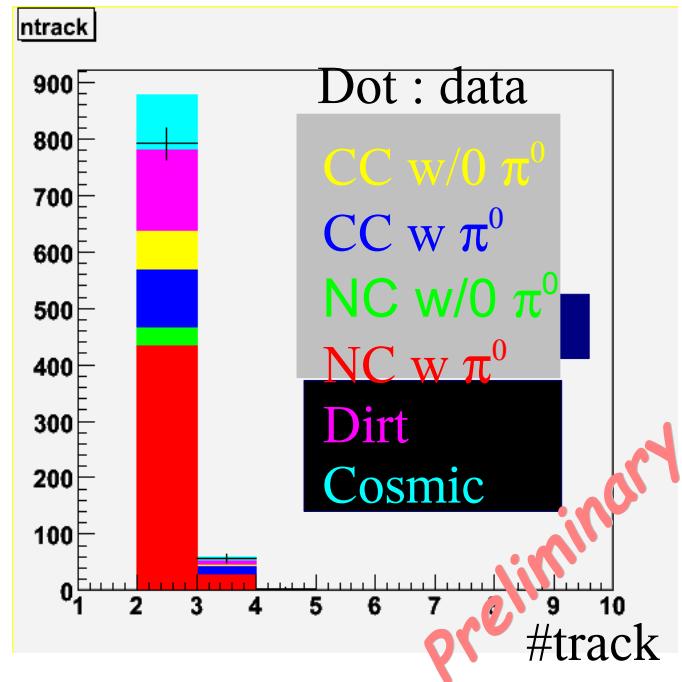
MC is normalized to  
SciBar-MRD matched sample

Fit  $P_\mu$ ,  $\theta_\mu$  distribution  
to get neutrino spectrum  
for  $\nu_\mu$  disappearance

Result will come soon

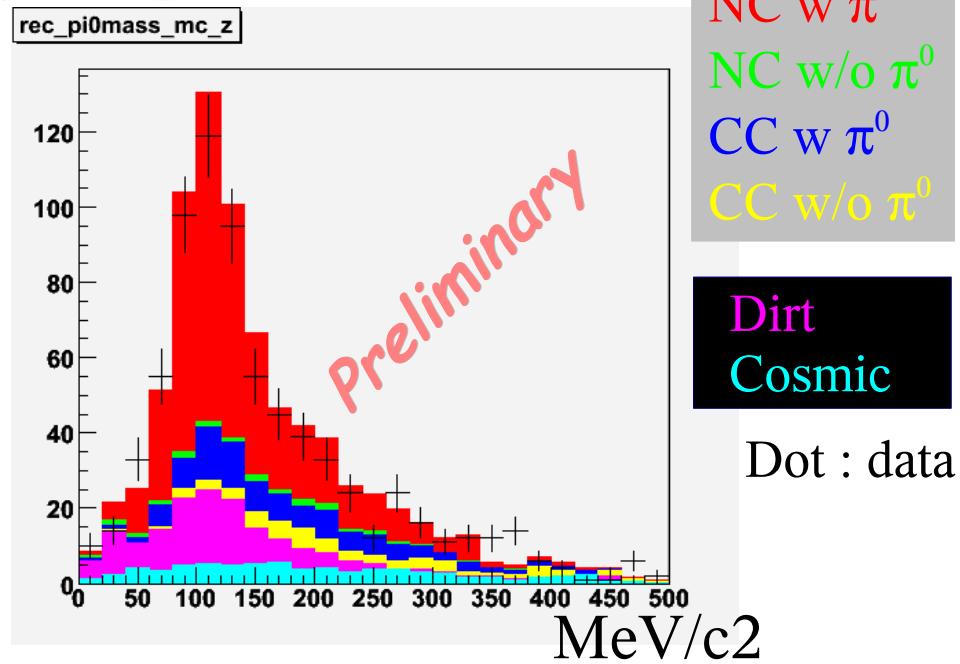
# Track Multiplicity

## Track Multiplicity after selection



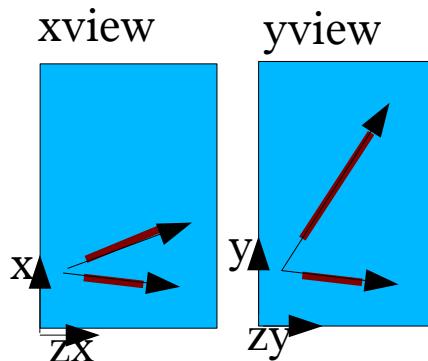
# Pi0 Mass (Detail)

Only one pair in a events

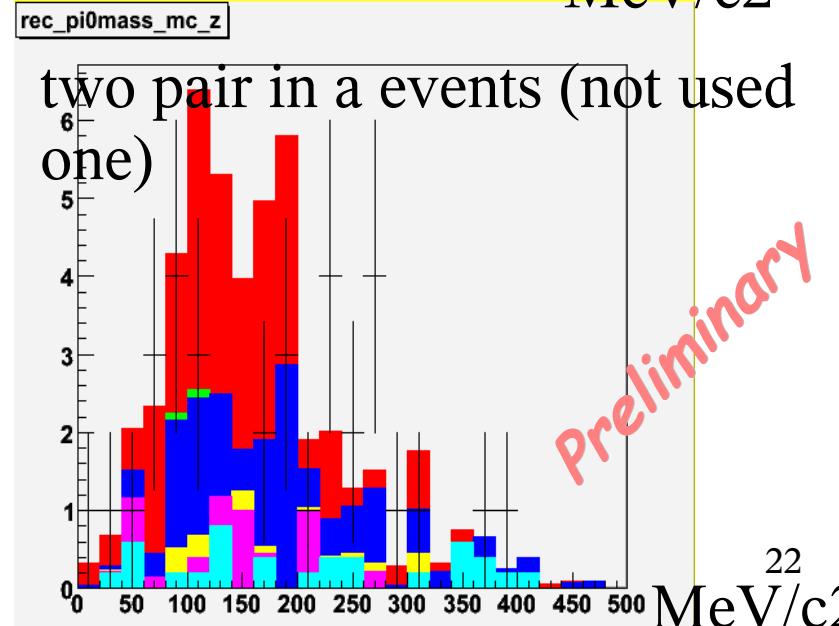
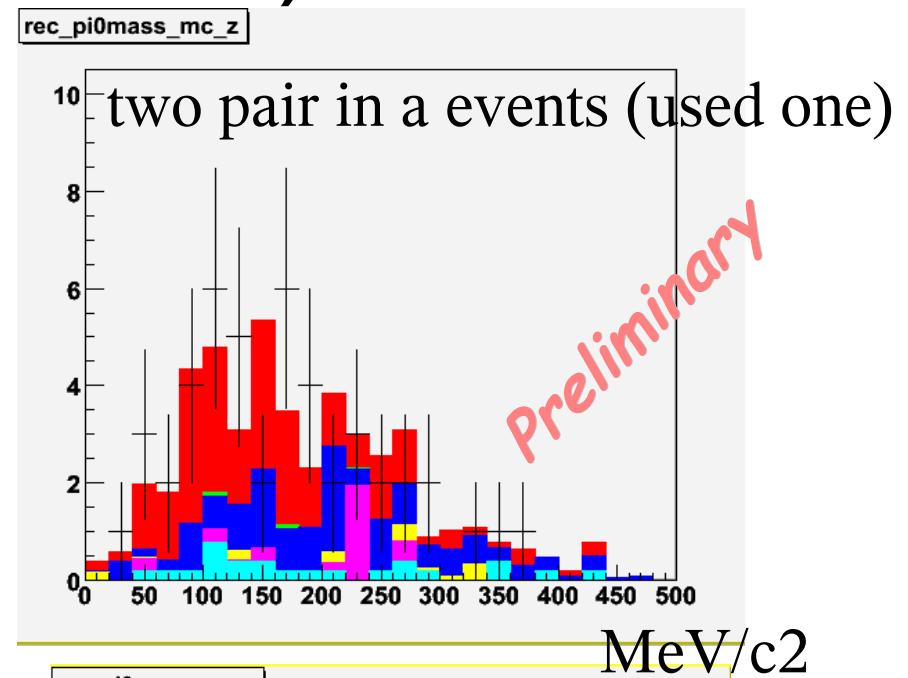


choose a pair with minimum  $|zx-zy|$

$zx, zy$ : the z position of the intersection of extrapolated tracks for each view

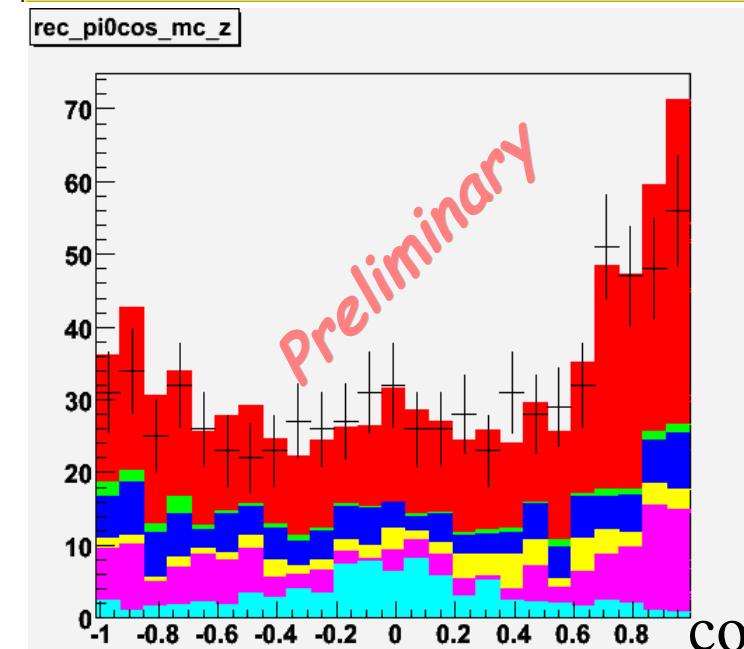
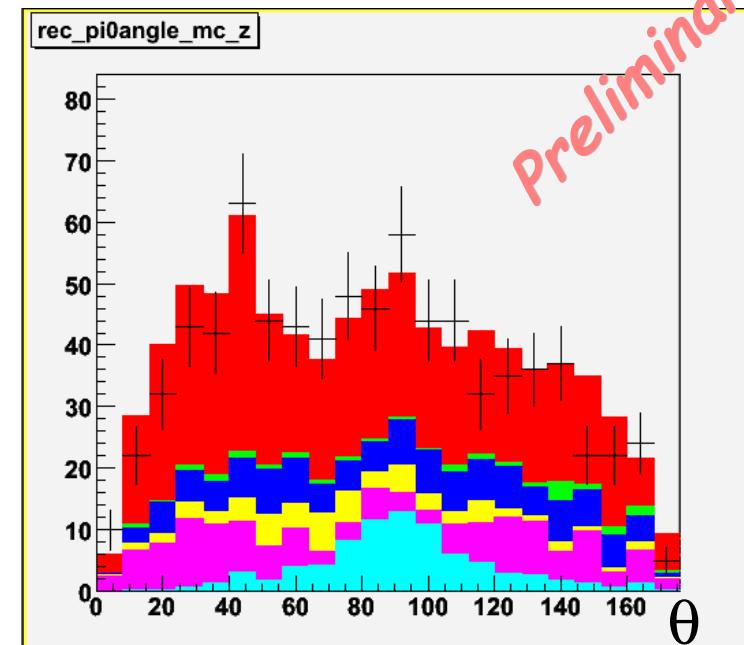
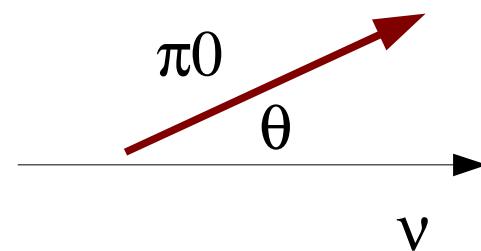
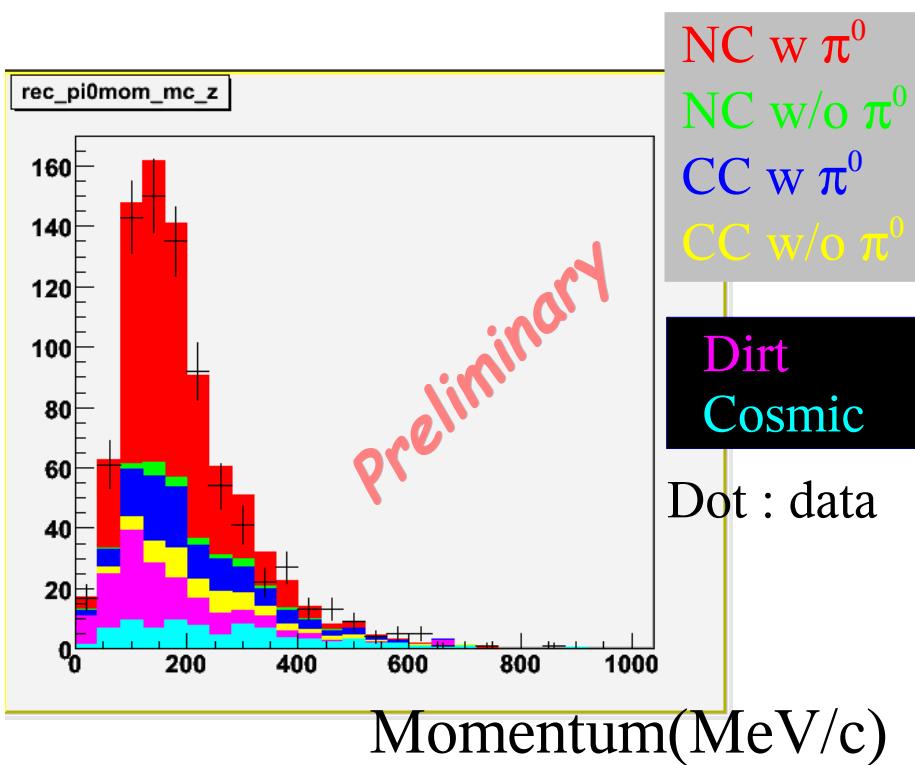


z: beam direction  
y: vertical (to sky)  
x: horizontal



22 MeV/c<sup>2</sup>

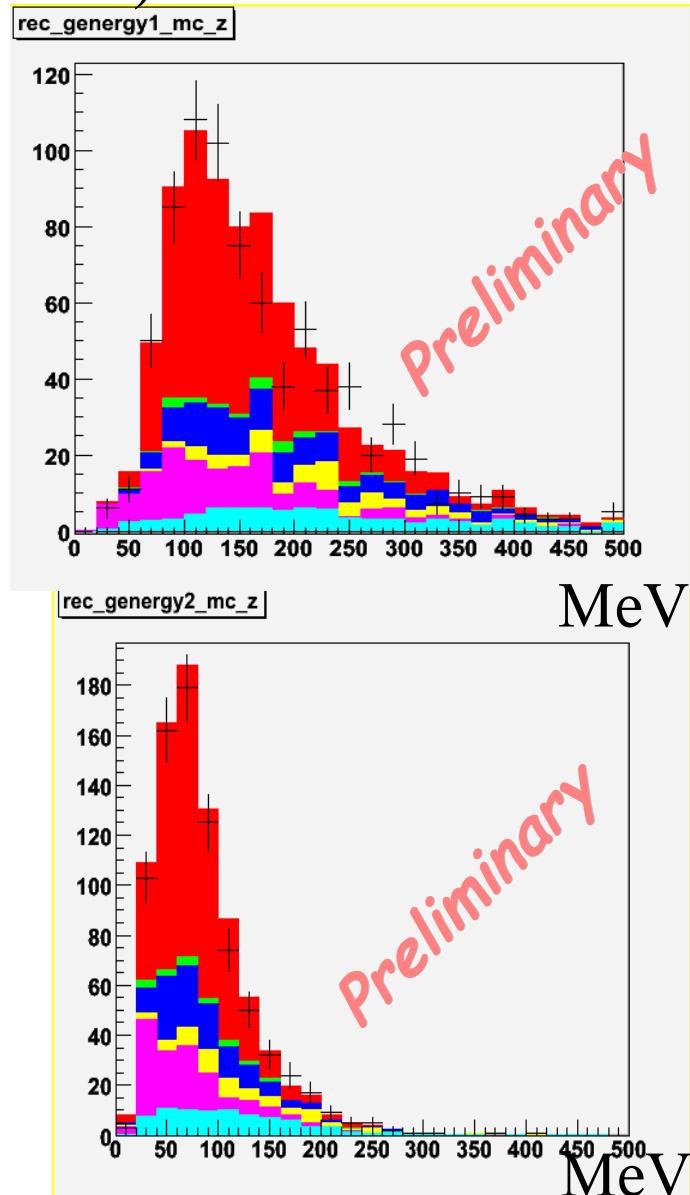
# $\pi^0$ Kinematics



# Gamma Kinematics

Energy

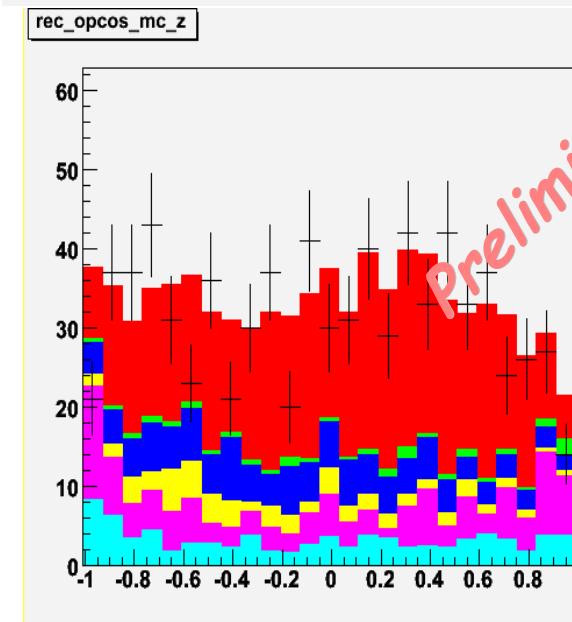
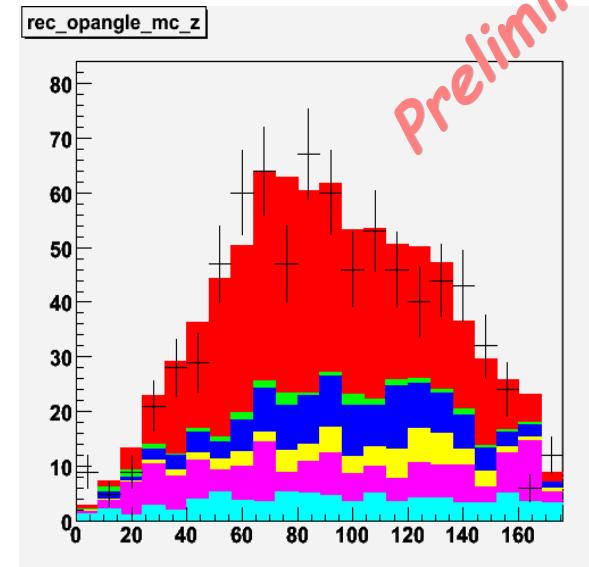
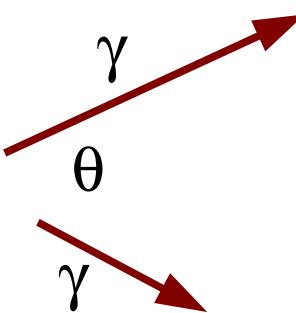
E1,E2 (E1>E2)



NC w  $\pi^0$   
NC w/o  $\pi^0$   
CC w  $\pi^0$   
CC w/o  $\pi^0$

Dirt  
Cosmic

Dot : data



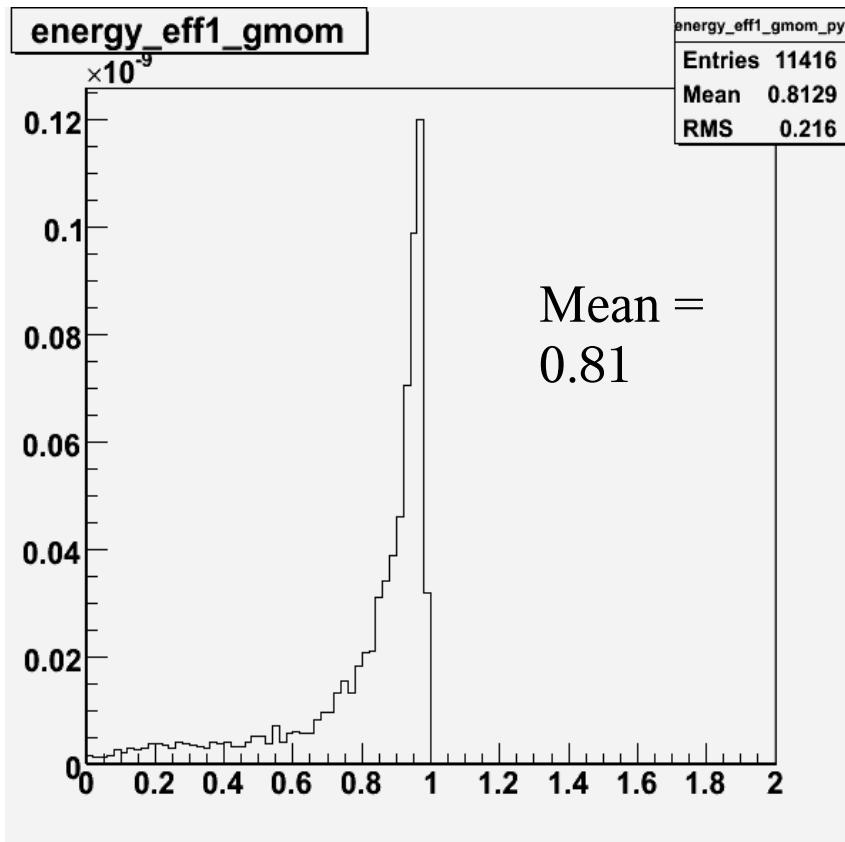
θ

cosθ

# $\gamma$ reconstruction performance using Extended Track (1)

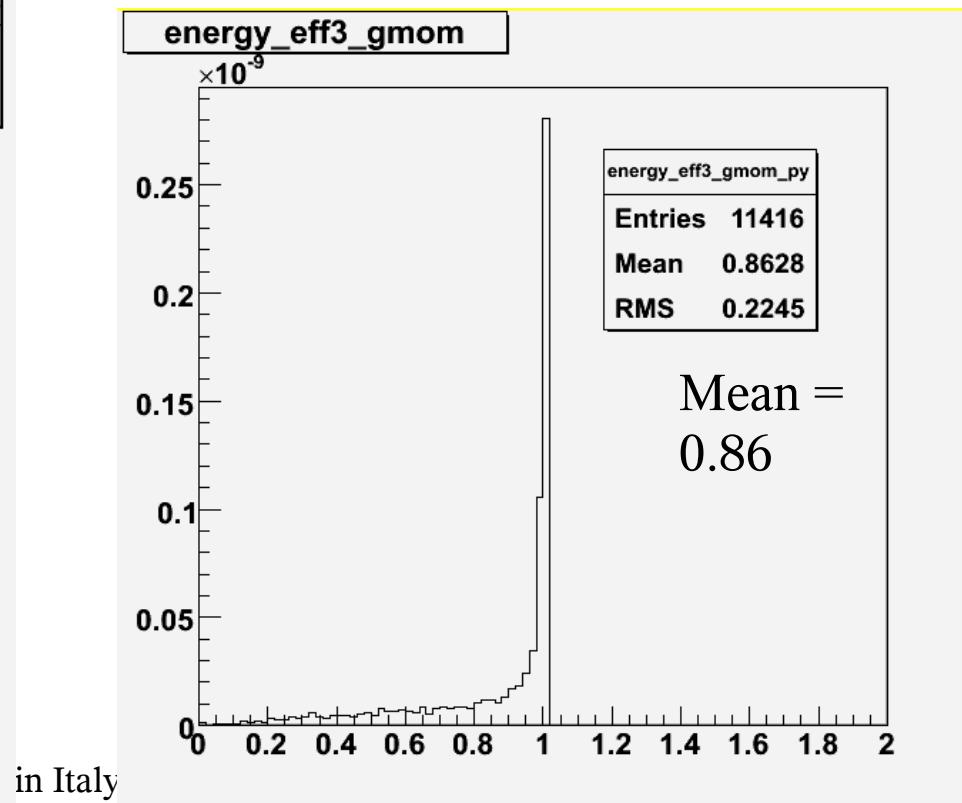
Energy Collection Efficiency

$$= \frac{\text{Energy due to } \gamma \text{ in ETrack}}{\gamma \text{ energy loss in SB}}$$



Purity of  $\gamma$  hits

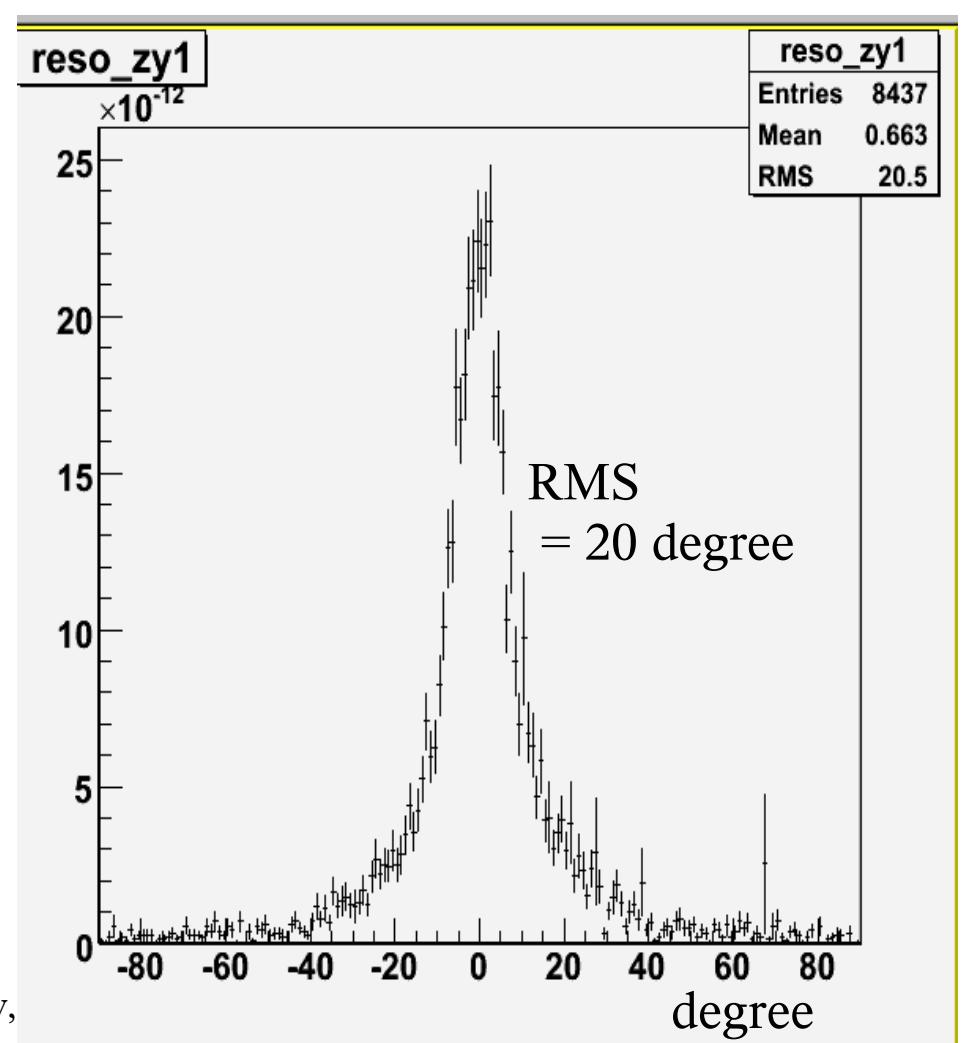
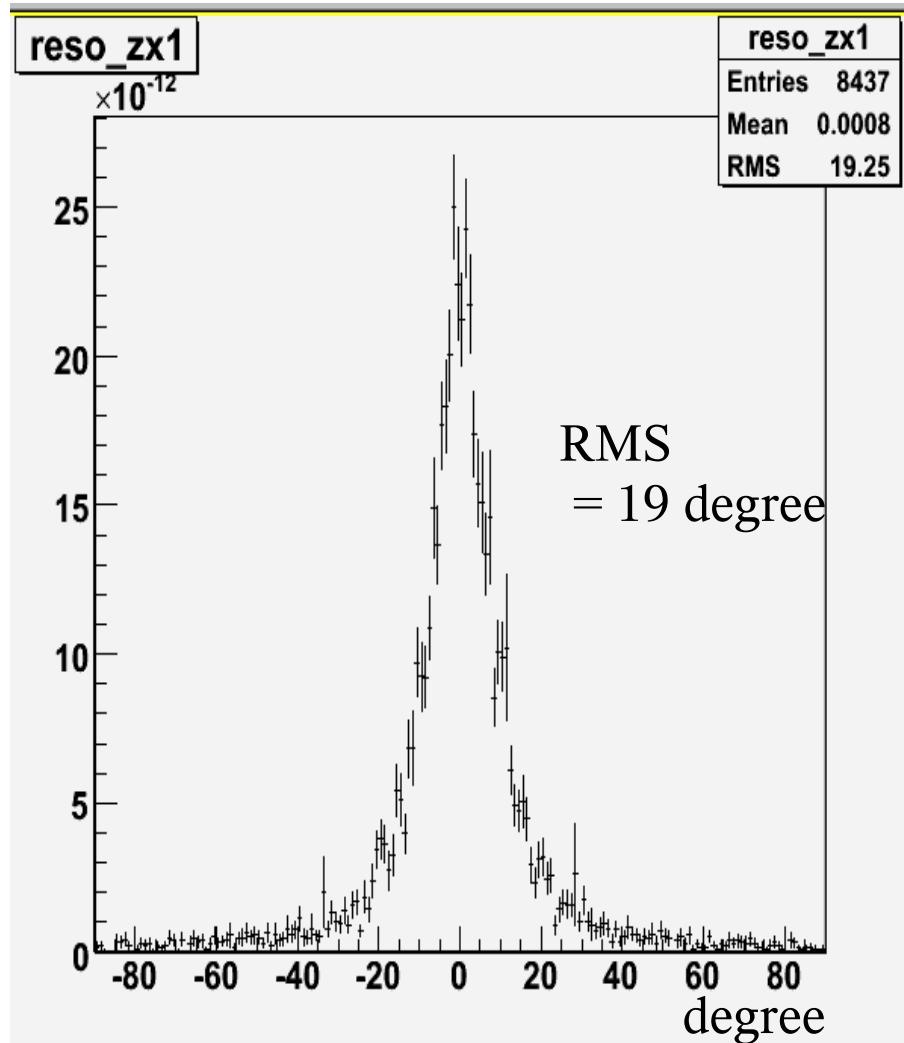
$$= \frac{\text{Energy due to } \gamma \text{ in ETrack}}{\text{Energy in ETrack}}$$



# $\gamma$ reconstruction performance using Extended Track (2)

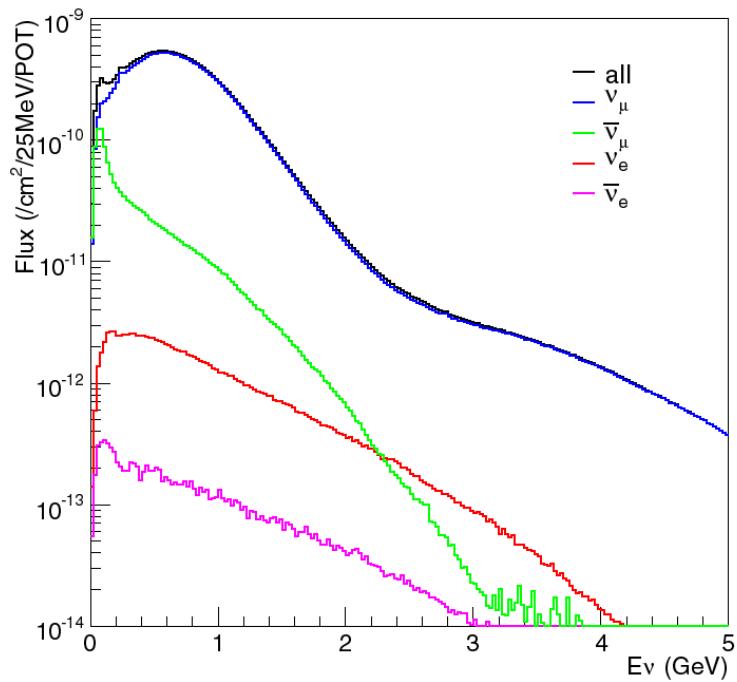
$\gamma$  angular resolution (x 2D)  $\theta_{xz}^{rec} - \theta_{xz}^{true}$

$\gamma$  angular resolution (x 2D)  $\theta_{yz}^{rec} - \theta_{yz}^{true}$

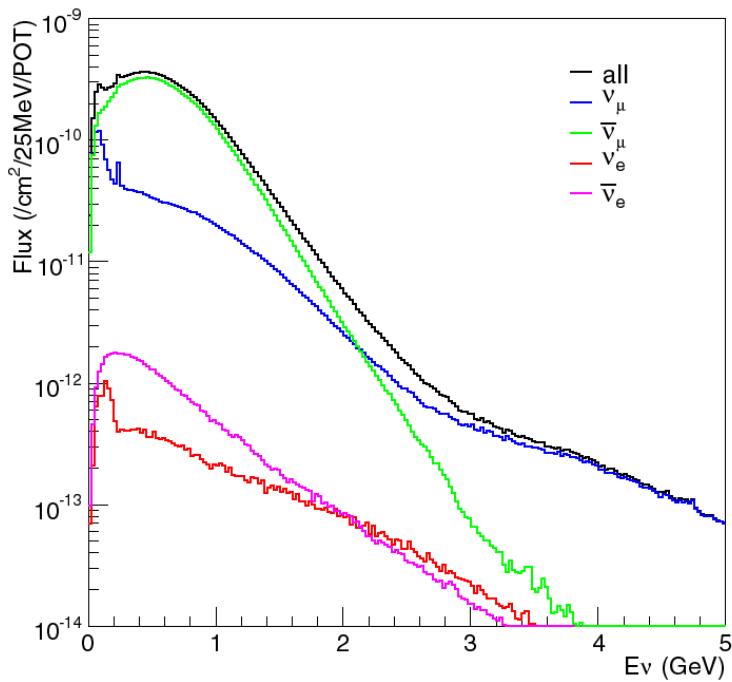


# Neutrino Flux at SciBooNE

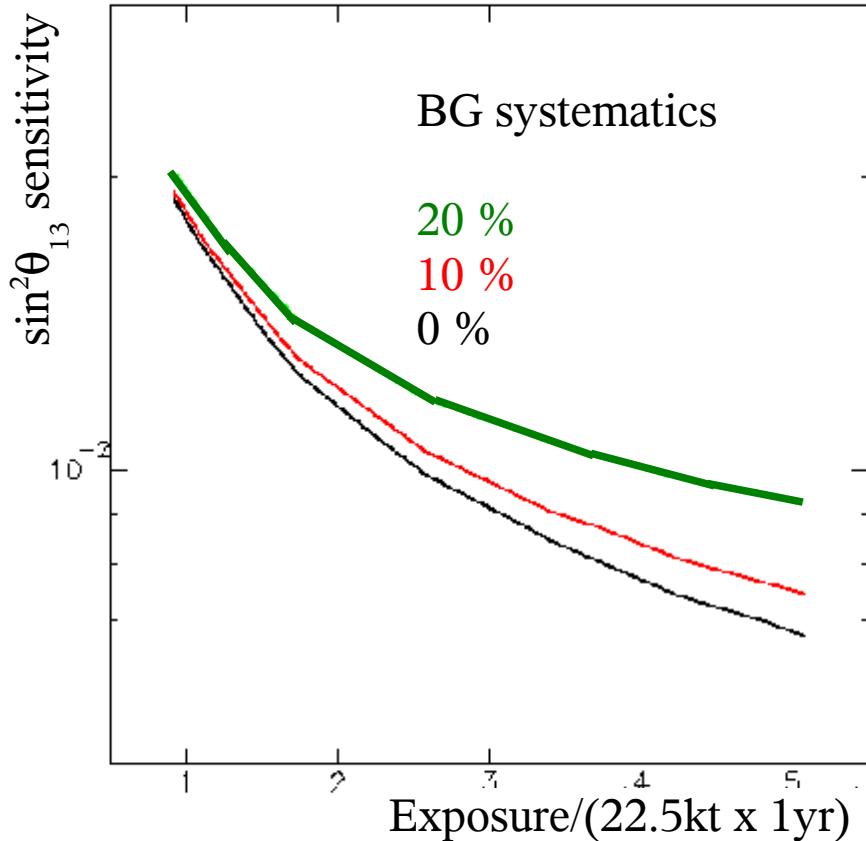
neutrino mode



antineutrino mode



# For T2K experiment



Oscillation probability

$$P(\nu_\mu \rightarrow \nu_e)$$

$$\sim \sin^2 2\theta_{13} \sin^2 \theta_{23} \\ \times \sin^2(1.27 \Delta m_{13}^2 L(km) / E(GeV))$$

- Want to reduce uncertainty in  $\sigma(NC\pi^0)$  from 20% to 10 %
  - improvement of factor of 2 in ultimate T2K sensitivity to  $\theta_{13}$
  - or 2.5 years vs. 4 years to  $10^{-2}$

# Event Summary

Data	NC w p.	BG (internal)	Dirt	Cosmic	efficiency(%)
Generated		120.64			
Pre-Selection	4228	1221	2389	718	179
p rejection	2102	781	953	507	154
$\mu$ rejection	1572	719	361	485	137
disconnection	1357	664	252	463	116
Trk merge	852	462	220	151	107